## REVIEW ON THE ITALIAN RADIO TELESCOPE RECEIVERS

## RECOMMENDATIONS

P. Bolli (INAF-OAA) (on behalf of the WG)

## AGENDA

- Background
  - Objectives (from #3 of ToR)
  - Deliverable (from #5 of ToR)
  - Process of the WG for elaborating the rec.
- Recommendations (incl. timeline and financial aspects)
  - on under development receivers
  - on future receivers
  - on management of the receiver group

## **OBJECTIVES 1/2**







## DELIVERABLE



## WHERE DO REC. COME FROM?



## HOW DO THEY LOOK?



# Recommendations

## **TWO PERIODS**

### 2017-2018

### 2019 and beyond

Almost all the receivers under construction should be completed

Development of future state-of-the-art receivers

# Either as it is currently planne or downgraded to 7 feeds

#### RECEIVER RT

SRT	S-band 3-4.5 GHz	Almost unique in the international scenario High-level science topics can be addressed with SD observations
SRT	Clow-band 4.2-5.6 GHz	Interest for VLBI (high publication rate) Hottest topics in today's radio astronomy
SRT	Q-band 33-50 GHz	Good opportunity to start testing metrology at relatively high frequency keeping at the same time a receiver with a high scientific impact No similar multi-feeds are available at foreign radio telescopes
SRT	ALMA 2-3 band 67-116 GHz	Time schedule is compatible with the SRT schedule. The ALMA receiver has remarkable high interest from a scientific point of view and allows for advanced technical performances The INAF-IASF group can take the responsibility of this task, thus there should be no additional workload on the receiver group.

## 2017-2018 for MED and NOTO

### RT RECEIVER

### WHY?

MED	Ku-band 13.5-18 GHz	Continuous coverage from 12 to 26 GHz. Interest for continuum studies & spectral line analysis. Relevant in the International context (like long term monitoring programs and expansion of the EVN capabilities in this frequency range)
ΝΟΤΟ	S/X/L-band 1.3-1.8 / 2.20-2.36 / 8.18-8.98 GHz	High scientific interest especially for VLBI and IVS observations.
ΝΟΤΟ	05.545-00.545 0112	Great interest for the participation to the observations of the GMVA network. Test the suitability of Noto (both telescope and site) for such frequency.
	would	the very limited cost of the adaptation of -MPIfR receiver for the secondary focus, we recommend to proceed with this solution. later stage, to adapt the ex-IRAM W-band rer.

# A full use of MED in the Q and s requires an antenna

### RT RECEIVER

Multi-feed in

W-band

Simultaneous

frequency in

**K**/Q/W-bands

SRT

SRT

MED

NOTO

PAF in C-band Relevant interest both as a technological demonstrator and as a new receiver to perform cutting-edge science (3 proposals). The interest for a PAF receiver matches with the involvement of INAF in the SKA AIP project (PHAROS2). We encourage to develop a new PAF in the C-band with state-of-the-art performances.

Considerable interest in such front-ends emerged from the Call for Ideas. Financial effort is being made by INAF to restore the active surface system of SRT to its best performances, opacity conditions at the SRT site are acceptable and RFI is not expected to be a concern in the coming years at such frequencies.

This front-end gives represents a niche in which also the smaller Italian radio telescopes can give a substantial contribution. Strong collaborations are already in place with the KVN and VERA arrays, and could be reinforced by adding MED (and possibly NOTO) to simultaneous high-frequency experiments

It needs to perform at nominal technical capabilities with also full operability of the frequency agility as well as to reach stable operational procedures

## **PROJECTS for FUTURE EVALUATION**

## RT RECEIVER

### WHY?

SRT / MED / NOTO	BRAND 1.5-15.5 GHz	<ul> <li>It is a new generation ultra-wide band receiver for VLBI, suited also for geodetic studies.</li> <li>This project pushes for advanced digital acquisition system especially to handle RFI issues.</li> <li>BRAND at SRT is a good opportunity in particular to perform geodetic studies since a classical S/X receiver is not available.</li> <li>Critical issues: 21 cm not available, high cross polarization, SRT primary focus crowded, able to satisfy the requirements for SD observations?</li> <li>The actual interest of BRAND will depend on its final design and capabilities.</li> <li>BRAND could be an interesting possibility also for the other two 32m radio telescopes.</li> </ul>
SRT	Camera in W-band 80-100 GHz	The bolometer proposed by de Bernardis is a very challenging project promising to increase the scientific applications of SRT and to widen the astronomical community interested in the use of the Italian radio astronomical facilities. A similar project has been recently developed at GBT. It should be developed under the responsibility of the proposing group. Its integration at SRT, as proposed, seems challenging due to dimensional and mechanical constraints and would need a close interaction with the SRT staff during the design phase.

# TIMELINE

Status	RT	RECEIVER	<b>'17 '18</b>		'19 and beyond
(0	SRT	S-band	OAC	OAC	
receiver:	SRT	Clow-band	IRA OAC OAA		The construction of these receivers is in a
ment	SRT	Q-band	IRA OAC	IRA	quite advanced stage and the workload is
Under development receivers	SRT	ALMA 2-3 band	IASF	IASF	distributed in the
	MED	Ku-band	IRA OAA	IRA	receiver group
	NOTO	S/X/L-band	IRA		
	ΝΟΤΟ	W-band		IRA	
New receivers	SRT	PAF in C-band	For each project, minimize the number of group to be involved		OAC
	SRT	Multi-feed in W-band			OAC
	MED	Sim. freq. in K/Q/W-bands			IRA

## FINANCIAL CONSIDERATIONS (k€)

Status	RT	RECEIVER	<b>'17-'18</b>	'19 & beyond
ers	SRT S-band		Fully funded	
SRT SRT SRT		Clow-band	Fully funded	
Under development	SRT	Q-band	600 (19 feeds) 180 (7 feeds)	
lopr	SRT	ALMA 2-3 band	80	
deve	MED	Ku-band	Fully funded	
der	NOTO	S/X/L-band	80	
nu N	NOTO	W-band	Negligible	
New receivers	SRT	PAF in C-band		~2700
	SRT	Multi-feed in W-band		~1700
	MED	Sim. freq. in K/Q/W-bands		~3000 (w AS) ~2200 (w/o AS)
TOTAL			760 (19 feeds) 340 (7 feeds)	~7400 (w AS) ~6600 (w/o AS)

## MAJOR UPGRADES ON EXISTING RX

### Instantaneous bandwidth

To enlarge the instantaneous bandwidth in the K-band receivers up to the whole band available. This will be done providing sub-bands 1 GHz wide by using new down conversion boards.

### **Low Noise Amplifiers**

The K-band receiver shows a bad reliability in some cryogenic LNAs, recurrently repaired. Today a substitution of all fourteen amplifiers is possible commercially, with a considerable mechanical effort inside the dewar in order to fit the different dimension of the LNA chassis.



## BEST PRACTISES FOR THE FUTURE

### Maintaining a leading position in receiver development within the International context

Project Scientist &	Local responsibility	Coordination	Call for Projects	Permanent commission
System engineering methodologies	Each project should be in charge of a specific group and	The development of radio astronomical Instrumentation	Periodically, a survey of the interest of the astronomical	Composed by astronomers and
Each new receiver should be assigned to a well-designed	locally managed	coordinated under the supervision of Section II of the INAF	community in new instrumentation should be conducted	technologists who will regularly meet to review the status of
management scheme to limit delays & financial	scientific collaborations among classical groups and	Scientific Directorate	by INAF	the ongoing projects and issue recommendations.
issues and to maintain high the scientific interest of the receiver	also extended to other INAF groups (IASF, SKA group) are encouraged			

# SPACE SCIENCE at SRT

- Recommendations regarding future ASI development of receivers dedicated to space science activities are out of the scope of this working group.
- However, we would like to point out the existence of an idea for a future receiver (Paolo Tortora's idea in X/Ka) that could be of interest for ASI.
- A critical issue is related to the RFI generated by ASI (receiving) devices. This aspect should be seriously considered keeping in mind the extremely sensitive radio astronomical receivers.
- The most relevant compatibility issue is related to the installation of high-power transmitters, whose installation needs a very detailed and accurate analysis to prevent damages to the INAF receivers and equipment.

# NORTHERN CROSS

- The use of the NC in the next years will be focused more on space science applications than on classical radio astronomical studies. No specific interests raised from the call for ideas on exploiting the Northern Cross for astronomical purposes.
- The NC is a propriety of the University of Bologna which should be involved in any discussion on possible upgrades.
- Very likely a significant refurbishment of the NC, like for example increasing the frequency band or the sensitivity, could make it very interesting for the low-frequency astronomical community.
- Decisions on possible upgrades of the NC are not of pertinence of this WG.

## THANKS FOR YOUR ATTENTION